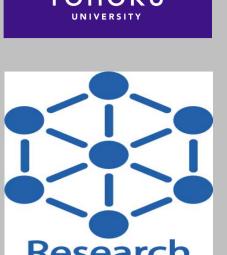
Vacancy formation energy in silicon studied by quenching under hydrogen atmosphere



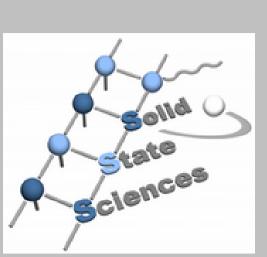


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Introduction

The vacancy formation energy in Floating zone (FZ)- and Czochralski (Cz)-grown Si crystals was determined by using a thermal quenching technique during which samples are saturated with hydrogen at the quenching temperature (1). The technique is based on measuring the optical absorption due to hydrogen atoms bound to vacancies in order to estimate the vacancy concentration. In high purity FZ Si, a vacancy formation energy of 3.85 to 4.0 eV was obtained, in excellent agreement with recent estimates by Bracht and Chroneos based on self- and metal diffsion experiments and ab initio calculations (2). At high quenching temperatures, the apparant vacancy formation energy in Cz-Si seems to be lower than in FZ-Si. The reasons for this difference are discussed and elucidated.

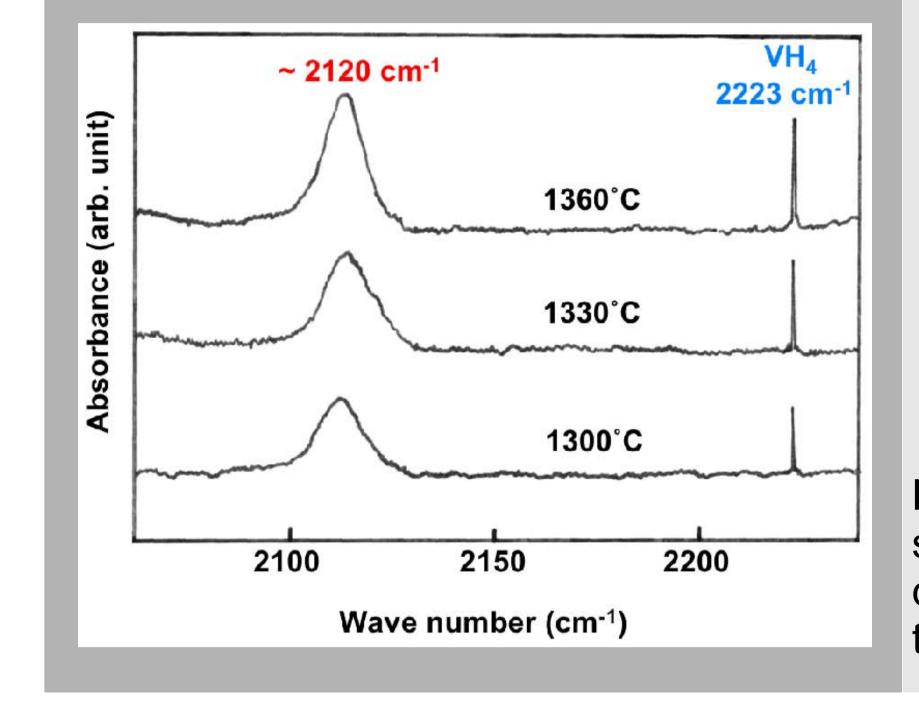


Fig. 1. FTIR absorption spectra of CZ Si samples quenched from different temperatures (3).

Experimental

Specimens were fabricated from a Cz-Si crystal with an oxygen concentration of 5.7×10¹⁷ cm³ as determined from the optical absorption coefficient at 9 µm assuming a calibration factor of 3.14×10¹⁷ cm². Specimens were cut with a diamond cutting machine, mechanically shaped with SiC powder and chemically etched with a mixture of HNO₃ and HF (5:1). The final specimen size was about 6×6×11 mm³. Specimens were sealed in quartz capsules together with hydrogen gas. The hydrogen gas pressure was adjusted to be about 1 atm at the quenching temperature. After heating for 1 h at the quenching temperature, the quartz capsules were dropped into water. Some specimens were annealed at 450°C for 1 h to transform the quenched-in vacancies into VH4, while others were isochronally annealed for 30 min at a range of temperatures with a 50°C interval. Optical absorption spectra were measured at about 7 K using a Fourier transform infrared (FTIR) spectrometer with a resolution of 0.25 cm¹.

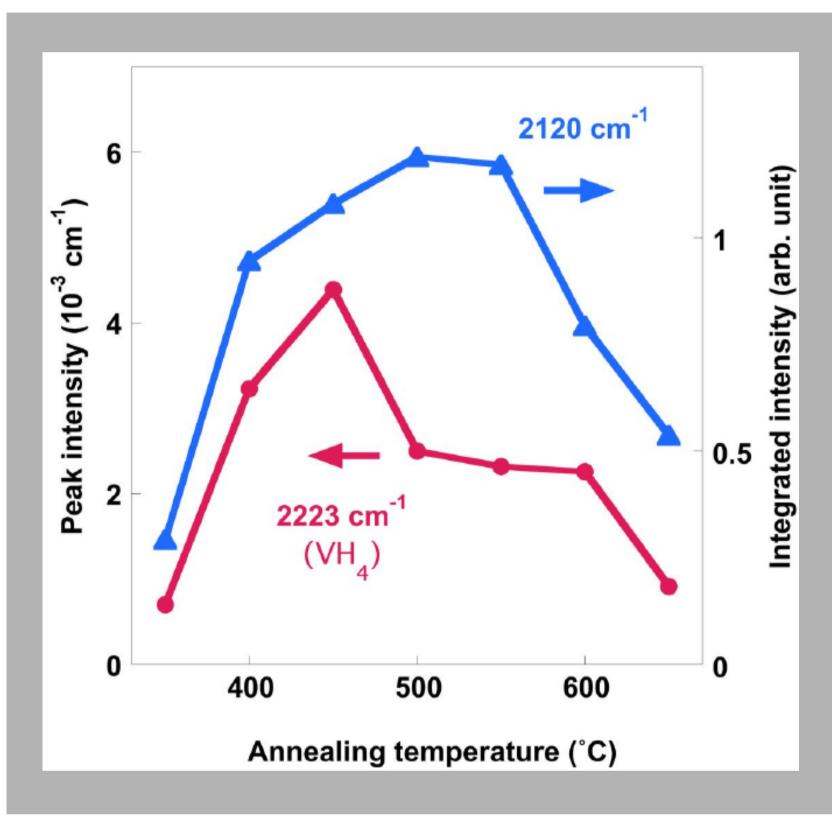


Fig. 2. Isochronal annealing temperature dependence of the peak intensity of the 2223 cm⁻¹ peak and the integrated intensity of 2120 cm⁻¹ peak of a Cz Si specimen. The specimen was quenched from 1364°C and then annealed during 30 min at each temperature.

Observations and discussion

As illustrated in Figs. 1, 2 and 3, in Cz-Si an additional absorption peak is observed at around 2120 cm⁻¹ different from the well-known 2223 cm⁻¹ peak that is associated with VH₄ (3). The defects responsible for the 2120 cm⁻¹ peak are proposed to be VOH₄ and VO₂H₄, whereby the VO and VO₂ concentrations are assumed to be determined according to the Voronkov and Falster model for grown-in defect formation in Cz-Si (4).

The formation energy of the 2120 cm⁻¹ related defect as estimated from the quenching temperature dependence of the integrated intensity of the absportion peak is about 2.3 eV. The small apparent formation energy of the vacancy in Cz Si at high quenching temperatures compared to that measured in high-purity FZ, is most probably due to the interaction between vacancies and oxygen atoms. It is assumed that oxygen atoms act as a source of vacancies which allows to explain the experimental observations (3).



$$C_{V} = (1 - ZC_{0}) \exp(-G/kT) + ZC_{0} \exp\{-(G_{0} - E_{R})/kT\},$$

with $C_{_{V}}$ the total vacancy concentration, Z the coordination number, $C_{_{0}}$ the impurity concentration, and $G_{_{0}}$ the Gibbs free energy for vacancy formation in the perfect crystal part. k is the Boltzmann constant. $E_{_{B}}$ is the interaction energy. For oxygen $ZC_{_{0}} << 1$ so that the total vacancy concentration increases with oxygen concentration (3).

After taking into account the contribution from vacancies generated in a perfect crystal, the formation energy of a vacancy at an oxygen atom was evaluated to be about 1.5 eV.

A similar reduction of the apparant vacancy formation energy and also a higher vacancy concentration than in FZ Si at high temperatures was observed in Ge doped Cz-Si (5).

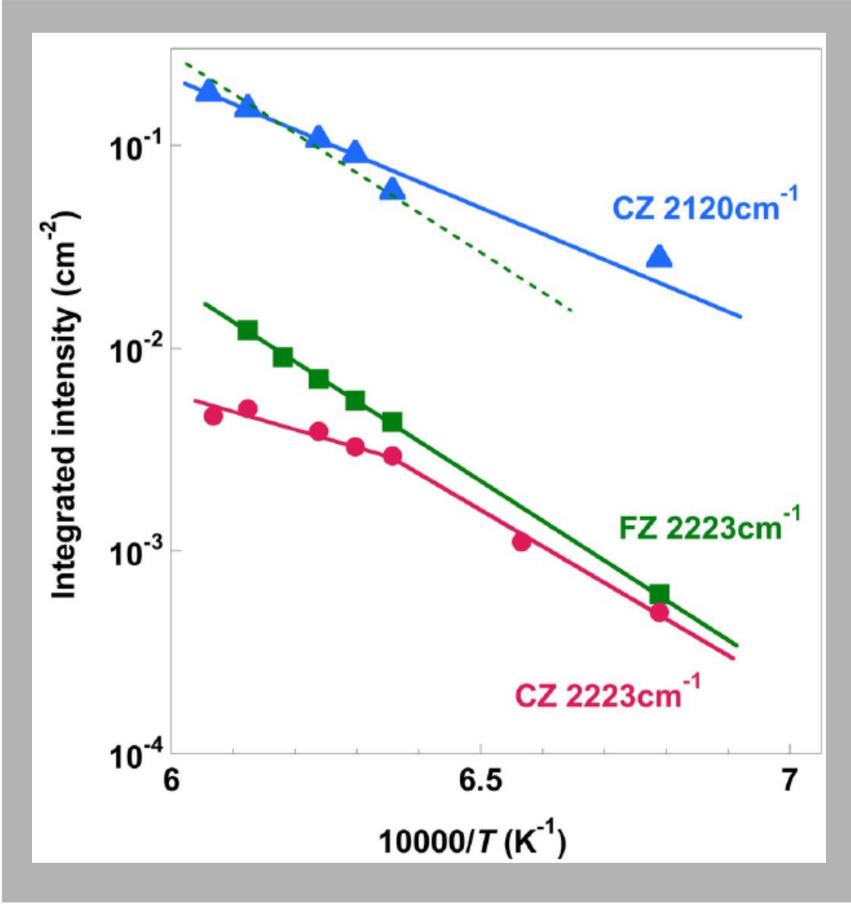


Fig. 3. Quenching temperature dependence of the integrated intensities of the 2223 and 2120 cm⁻¹ peaks.

Acknowledgments

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